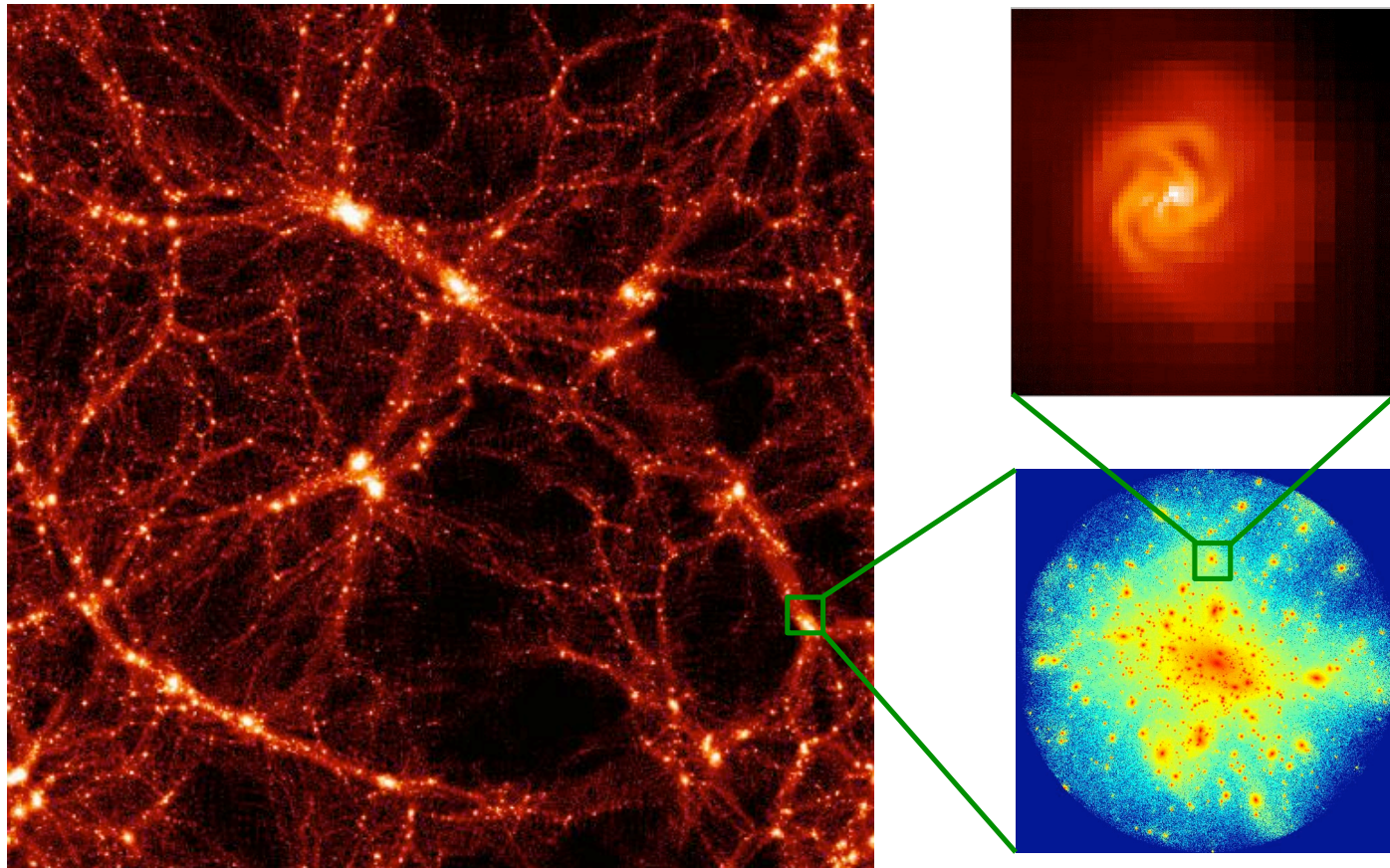


Extreme Computational Cosmology



- **Horizon: what is it ?**
- **Initial conditions**
- **Physics**
 - **From pure N body to fully reactive radiation MHD ?**
 - **Star formation and supernovae driven winds**
 - **AGN driven outflows**
- **Supercomputers in Europe. Current and future projects**
- **Goal of this workshop**

5 Partners

<u>Labo</u>	<u>Co-I</u>	<u>Lieu</u>
LUTH	J.-M. Alimi	Meudon
IAP	S. Colombi	Paris
LERMA	F. Combes	Paris
CRAL	B. Guiderdoni	Lyon
SAP	R. Teyssier PI	Saclay

Horizon Scientist: collaborator responsible of a work package under the supervision of a node leader
30 scientists and 10 students

Horizon Associate: collaborator without work package participating to Horizon on a short term basis.

Executive committee: 5 co-I meeting every month

Scientific committee: 10 members meeting every year (in preparation)

Horizon Web Site: with separate private and public parts (1 technical and 3 editorial supervisors)
<http://www.projet-horizon.fr>

Horizon Mini-Grid: 6 quad AMD64 16 Gb servers located in each lab and interconnected as a grid (3 system managers)

Horizon Meso-Machine: 3 quad AMD64 64 Gb servers located in HPC1 and 250 000 additional “on demand” hours

Supercomputing Centers: combined proposals in France (IDRIS, CINES et CCRT) and in Europe (DEISA initiative)
<http://www.deisa.org>

Centres de calcul



Méso-machine



Post-traitement et
Archivage lourds

Soumission de jobs

Visualisation,
Post-traitement et
Archivage légers

Mini-grille



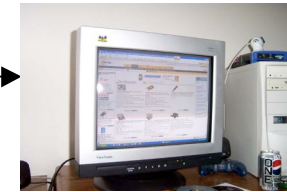
Paris



Meudon



Saclay



Lyon



Marseille

Fév. 2004: support from the French Galaxy and Cosmology Programs.

Avr. 2004: support from the French Astroparticle Program
Funding of the mini-grid 105 k€ (INSU + IN2P3)

Juin 2004: Support from Paris University (Paris 6)
100 k€ grants for travel and equipment up to 2008

Sep. 2004: Kick-Off Meeting: 1st Horizon Workshop

Fév. 2005: support from the French Astronomy Program (INSU)
Funding of the meso machine 120 k€ (INSU) + 30 k€ (CEA)

Sep. 2005: HP provides the meso machine as “on demand computing” HPC1

Oct. 2005: 2 post-docs are funded from CNRS (Saclay, Obs. Paris)

Oct. 2005: support from the French Science Foundation (ANR)
500 k€ grants up to 2008 (including 3 post-docs)

Oct. 2005: Horizon is selected by DEISA with 27 other European projects
Mare Nostrum computer in the Barcelona Supercomputing Centre

Nov. 2005: 2nd Horizon Workshop

3 box sizes: **500, 100 and 20 h^{-1} Mpc**

Unique set of initial conditions: **4096^3 , 2048^3 , 1024^3 , 512^3 ...**

3 types of simulations

- “periodic box”
- “zoom” on pre-identified halos
- “idealized” on pre-identified halos

Several types of models

- Pure N body + semi-analytics post-processing
- N body and gas dynamics “The Works”
- Isolated halo with prescribed boundary conditions

Several types of codes

- PM-AMR (RAMSES, ENZO, PMCOLL...)
- TREE-SPH (GADGET, MULTIZOOM...)
- ...

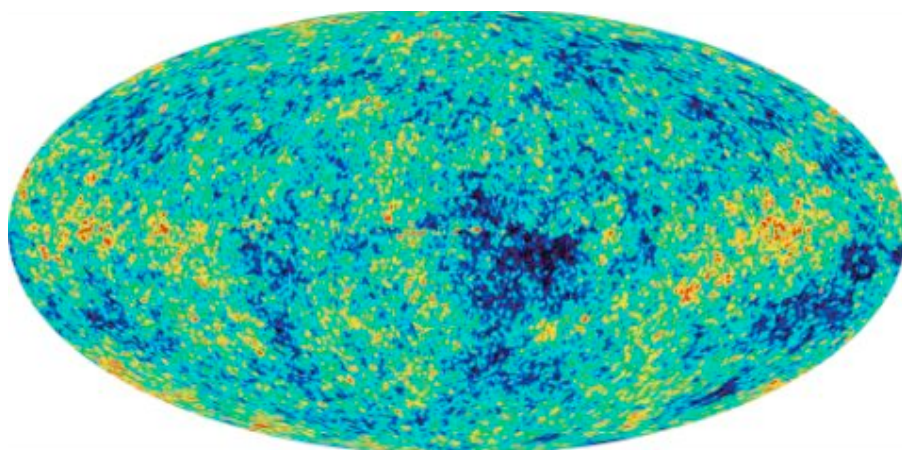
Several “on line” outputs

- halos (sub-halos) and galaxy catalogs, merging trees
- “all-sky” or “patch” virtual images (γ , X, visible, IR, mm, radio)
- mock spectra and spectro- images
- “raw data”, with images and movies

From current cosmological constraints, generate density and 3-velocity (3-displacement) fluctuations. Analytical transfer function with “baryons wiggles”. **2 fluids ?**

Standard procedure:

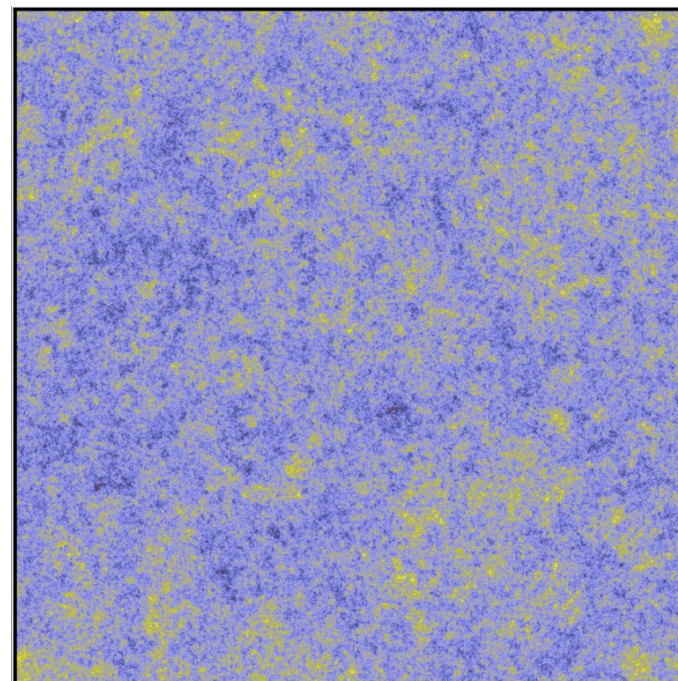
- generate white noise array in real space
- convolve by $\sqrt{P(k)}$ using FFTW



2D temperature map of the CMB observed by WMAP

Mare Nostrum IC from Yepes, Hoefft & Gottloeber
50 h^{-1} Mpc 2048^3 : 32 Gb per field

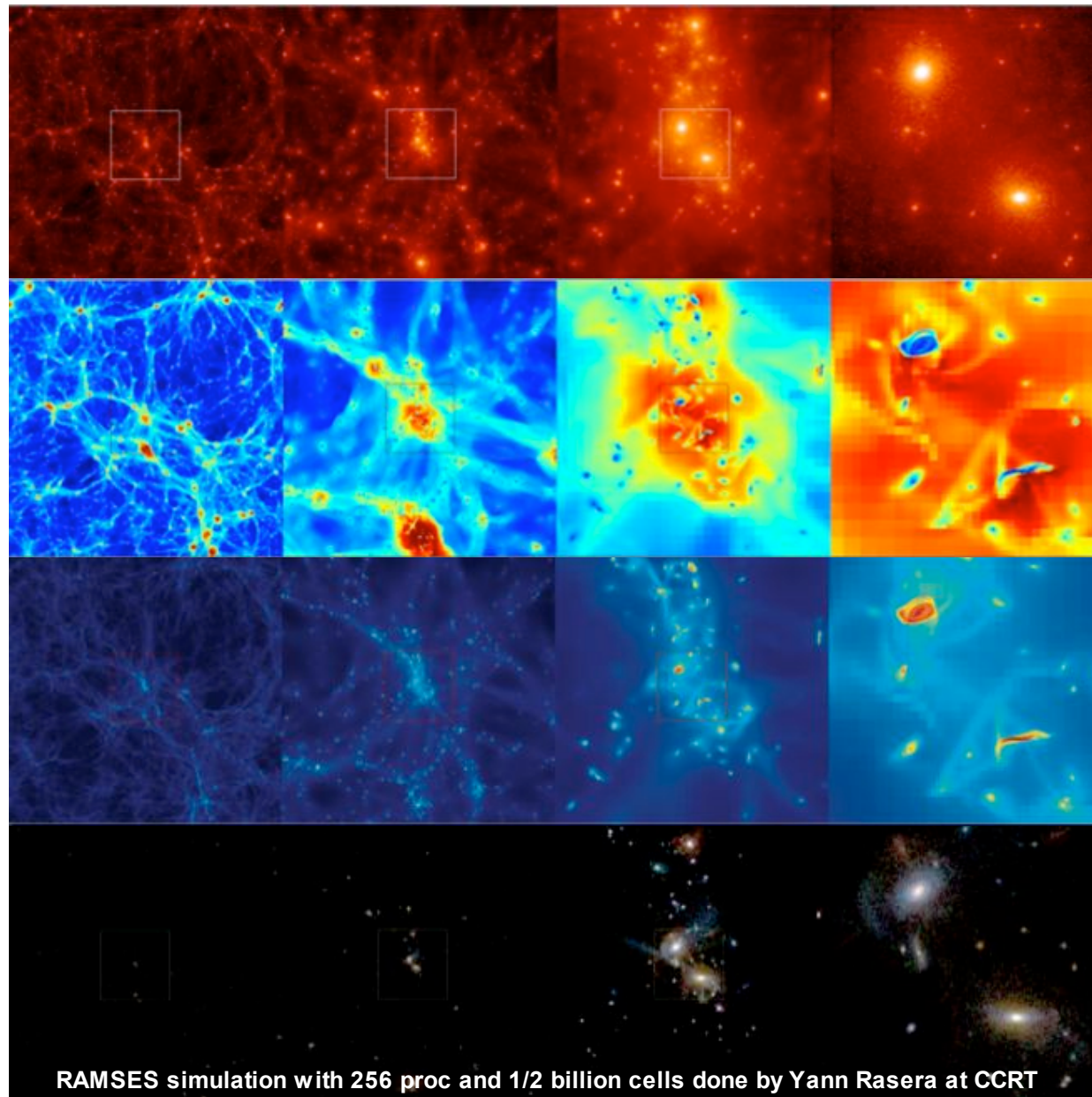
Horizon: 100 h^{-1} Mpc 4096^3 : 256 Gb per field
Same white noise for 500 and 20 h^{-1} Mpc



3D density field 2048^3 generated by MPgratic
Aubert, Pichon, Prunet

Zoom initial conditions will be
extracted from these 3 reference sets

Galaxy Formation in a 10 h^{-1} Mpc box @ $z=3$

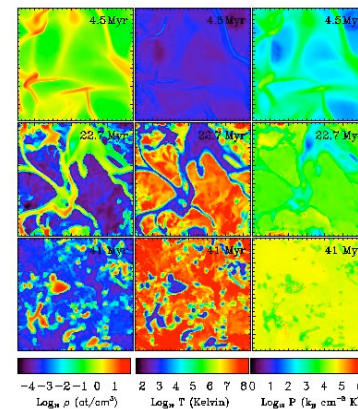


Metal dependant cooling and heating. Radiative transfer ?

Star formation within a multiphase/turbulent medium
 $t_* = t_0(\rho/\rho_0)^{-1/2}$ for $\rho > \rho_0$ with $t_0 = 1-10$ Gyr and $n_0 = 0.01-0.1 \text{ cm}^{-3}$
 Yepes et al., Springel & Hernquist

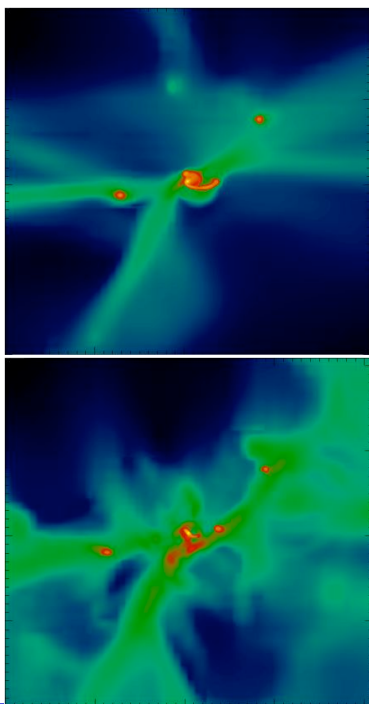
MHD ?

Supernovae heating: effective (polytropic) equation of state
 $P_* = P_0(\rho/\rho_0)^\gamma$ for $\rho > \rho_0$ with $T_0 = 1-5 \times 10^4 \text{ K}$. Starburst ?

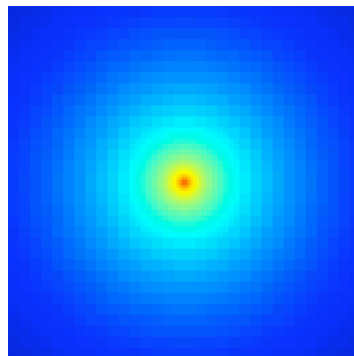


BGK simulation A. Slyz et al.

ENZO simulation
J. Devriendt

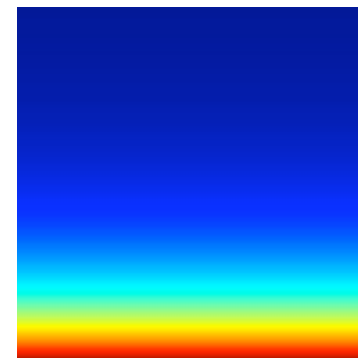


RAMSES simulation
Y. Dubois



AGN model and feedback ?

RAMSES simulation
A. Cattaneo



Galactic winds by supernovae kinetic feedback

- cooling delayed during $t_d = 10-100 \text{ Myr}$
- shrapnel (mass loading) deposited a few cells away

Superwinds ?

Clusters: Mare Nostrum (5th in Top500 list)

- . 4 Gb nodes with 2 processors (IBM blades)
- . Fast interconnect with Myrinet
- . up to 2400 nodes !

Blue Gene: LoFar (6th in Top500 list)

- . 2000 PowerPC 440 per rack and 6 racks (64 racks in the US !!)
- . ultra-fast interconnect (tore 3D + fat tree)

Technical issues and constraints

- . Parallel I/O and robustness
- . Load balancing
- . Data retrieval (internet access up to 20 Gb/s ?)

Prospectives

- . PITAC report: <http://www.nitrd.gov/pitac/reports>
- . Sartorius report: <http://www.recherche.gouv.fr/rapport/calcul/2005-017.pdf>
- . European “Extreme Computing Initiative” and beyond...

Horizon: a 3-year computational project to study large scale structure and galaxy formation

Extreme computing in Europe: <http://www.deisa.org> 27 projects selected
1/2 of them are astronomy and 1/4 of them are computational cosmology

Mare Nostrum: Yepes *et al.*

Mare Nostrum: Horizon

Millenium: White *et al.*

...

LoFar: Joop Shaye *et al.*

Coordination-comparison-competition ?

Goal of this workshop: *set up the run parameter file !*

What are the current *outstanding* questions in galaxy formation ?

What physics need to be simulated ? **Minimum and Goal**

What simulations need to be performed ? **Minimum and Goal**

Agenda:

1. Mare Nostrum run before summer 2006: 10^{10} AMR cells
2. Other runs to come (winter 2006 and beyond)

End